

Resolving the great drying of Mars: sequence stratigraphy of Aeolis Dorsa

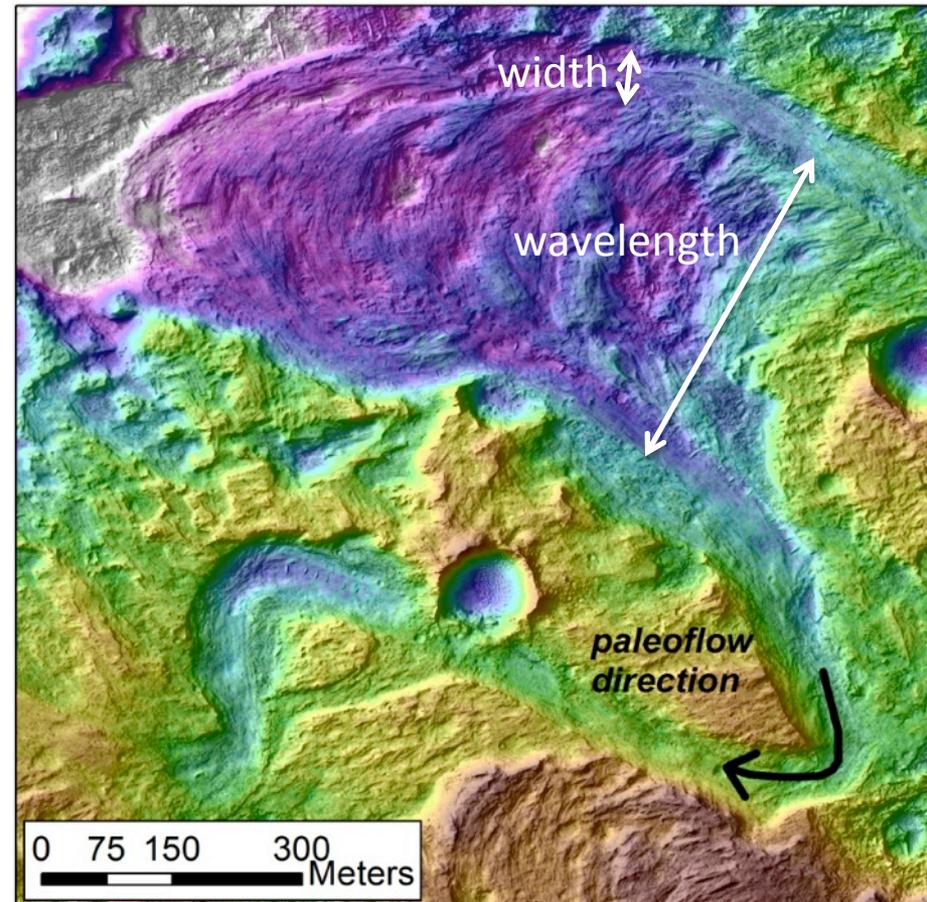
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Oded Aharonson (Caltech, Weizmann) & John Armstrong (Weber State)

Rationale: River-deposit dimensions constrain hydrology and climate on Early Mars, but stratigraphy is essential to build a time series of constraints on climate change

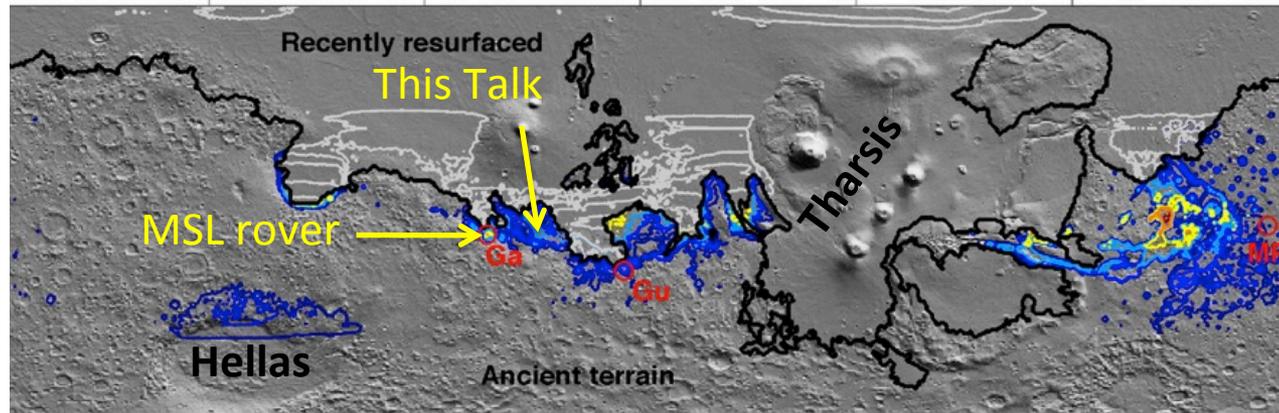
Today, use geologic mapping to:

- 1) Order river deposits by relative time
- 2) Determine if drying-out of Mars was steady or unsteady



Making best use of geologic proxies for paleoclimate in Aeolis Dorsa requires stratigraphy

Mars latitude



Mars longitude

Map of Early Mars water-availability model output (Kite et al., 'Seasonal melting ...' Icarus 2013a)

Aeolis Dorsa is a 10^5 km² low-latitude sedimentary-rock basin, $\sim 10^\circ$ E of MSL rover
- more river deposits than the rest of Mars put together

Parameters:

- Discharge $10^1 - 10^3$ m³/s
- Age Hesperian (or older)
- Deposition interval $>(1-20)$ Ma
- Atmospheric pressure $<760 \pm 70$ mbar (stat. error)

Burr et al., JGR 2010

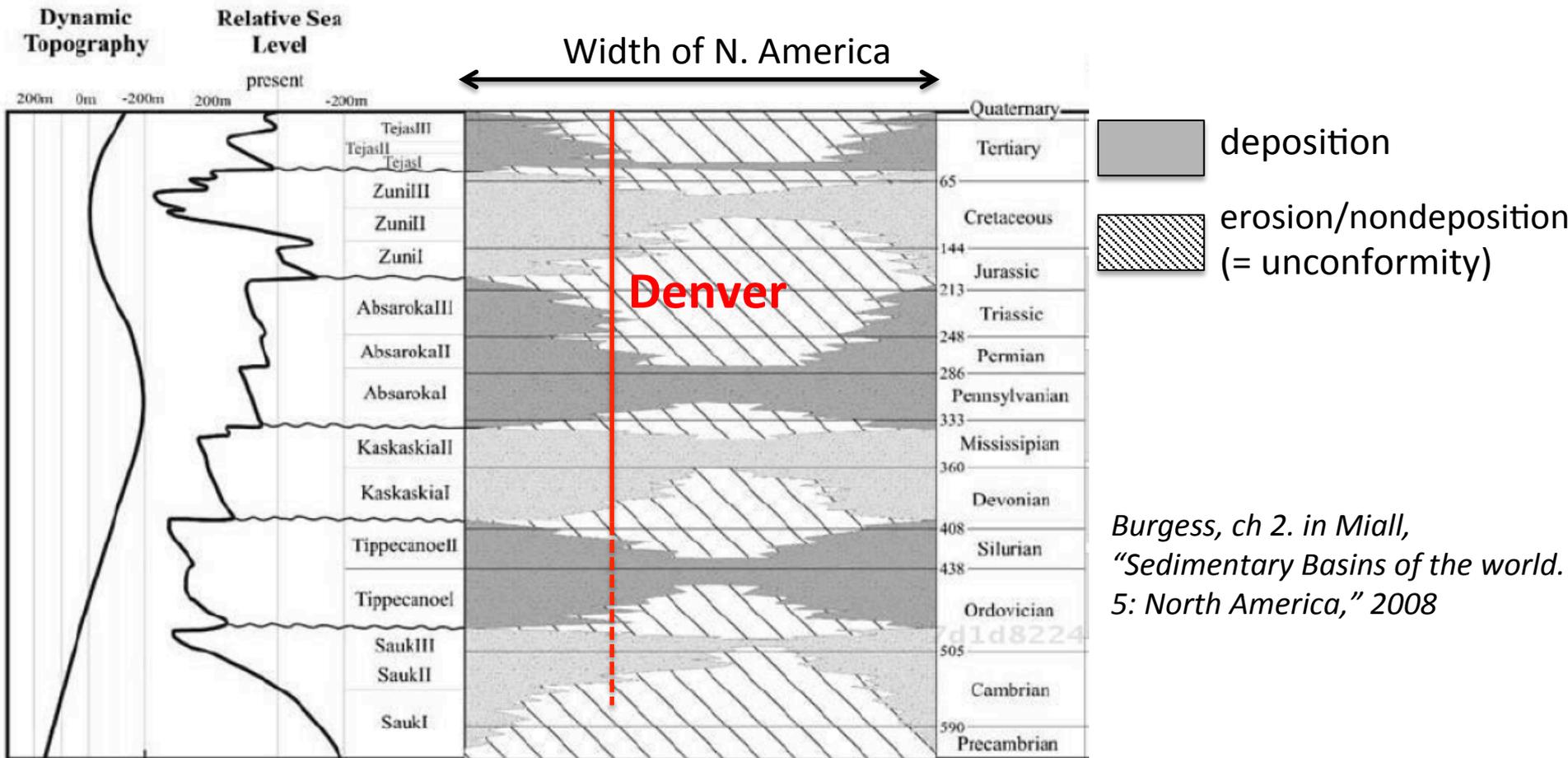
Zimelman & Scheidt, Science 2012

Kite et al., 'Pacing fluvial ..,' Icarus 2013

Kite et al., 'Paleopressure ...,' arXiv:1304.4043,
'accepted in principle,' Nature Geoscience

At Aeolis Dorsa we can constrain magnitude, duration, intermittency, and number of wet events, but we need to put river deposits in relative-time order to get a time series.

Key step in reading history of any sedimentary basin: identify unconformity-bounded sequences



Cause of alternation between deposition and erosion: **Global sea-level change** (plate tectonics), **dynamic topography** (plate tectonics)

On Mars, unconformities within layered strata appear uncommon

Edgett (Mars J. 2005), Wiseman et al. (JGR-Planets 2007),
 Milliken et al. (GRL 2010), Holt et al. (Nature 2010)

River deposits are eroding out of mappable geologic units

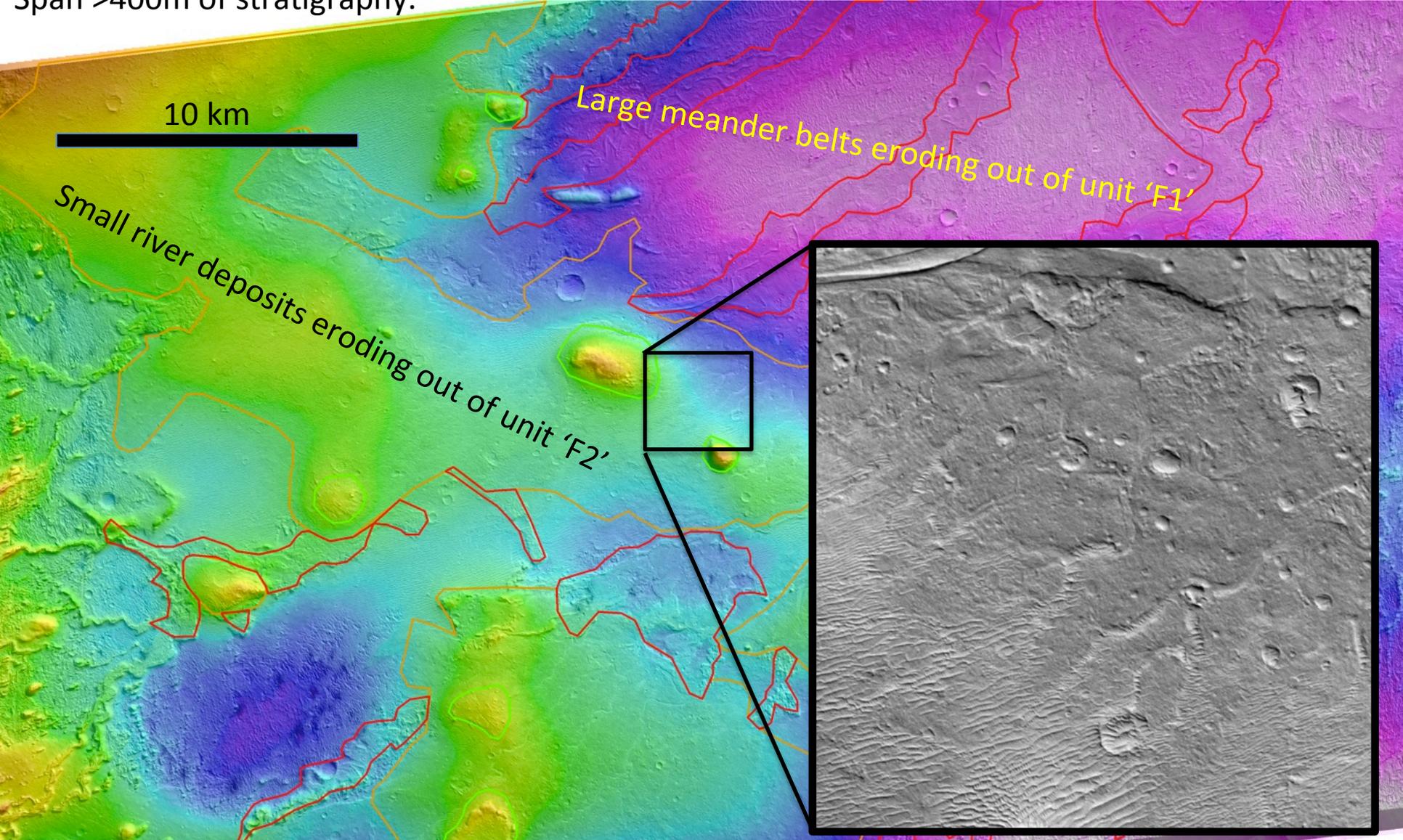
Span >400m of stratigraphy.



10 km

Small river deposits eroding out of unit 'F2'

Large meander belts eroding out of unit 'F1'



Area of B20_017548_1739_XI_06S206W

River deposits are eroding out of mappable geologic units

In time order:

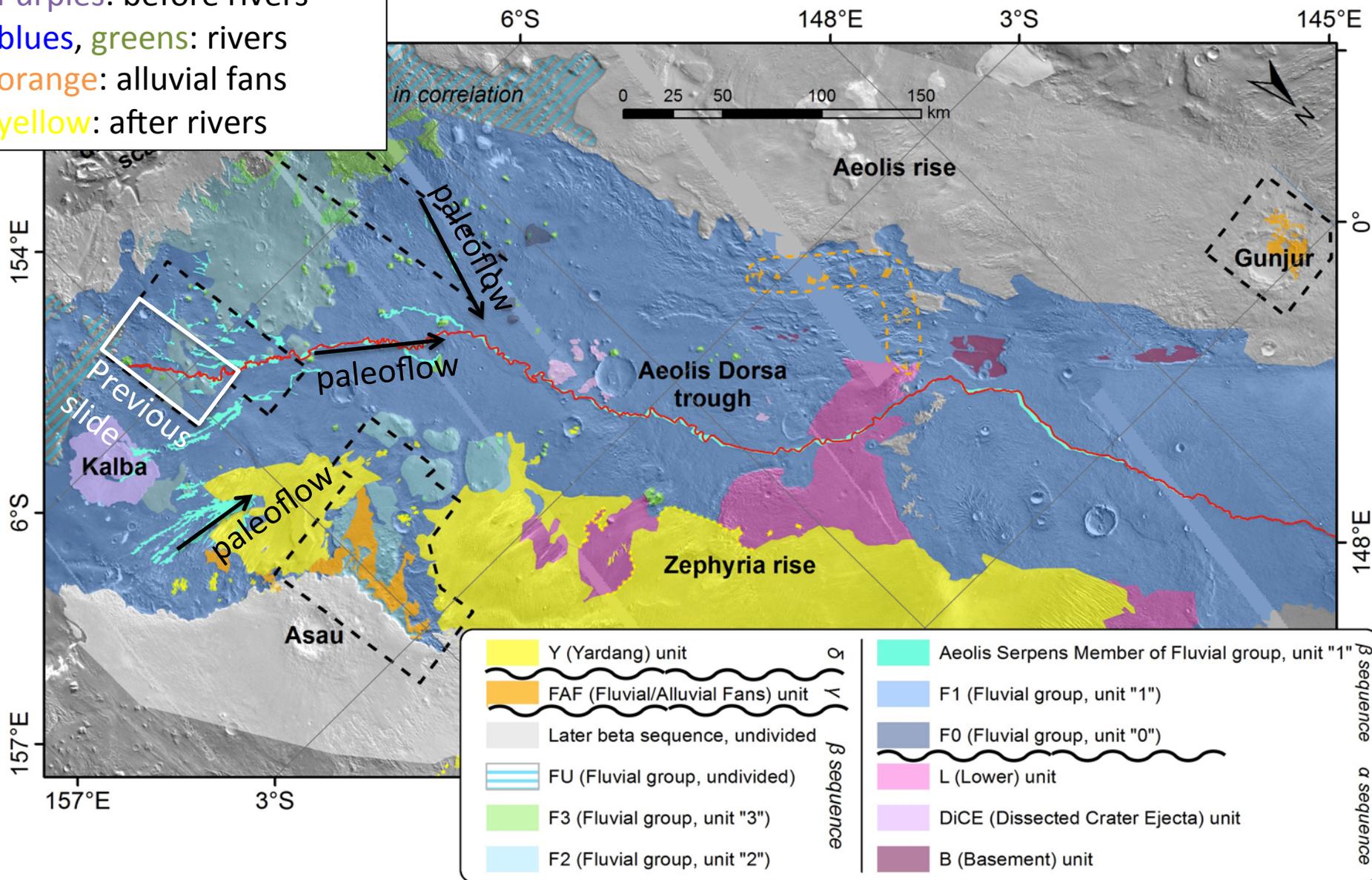
Purples: before rivers

blues, greens: rivers

orange: alluvial fans

yellow: after rivers

Coherent paleoflow direction ($n = 281$ meanders picked)



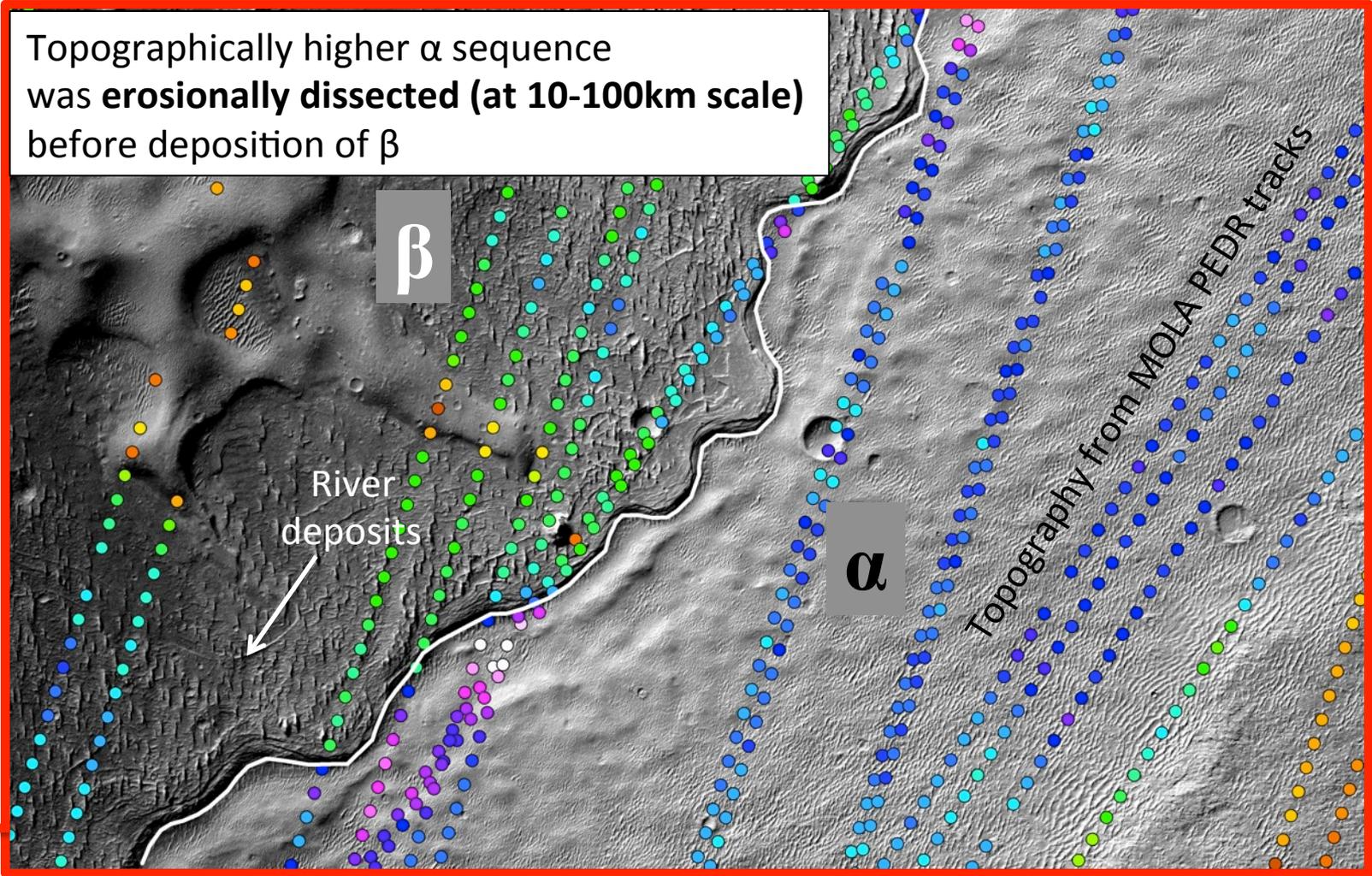
Sequence boundary example 1/3: **Pre-river layered sediments (α) were erosionally dissected before deposition of river sediments (β)**

δ

γ

β

α



Additional evidence for this unconformity: R.M.E. Williams et al., Icarus 2013

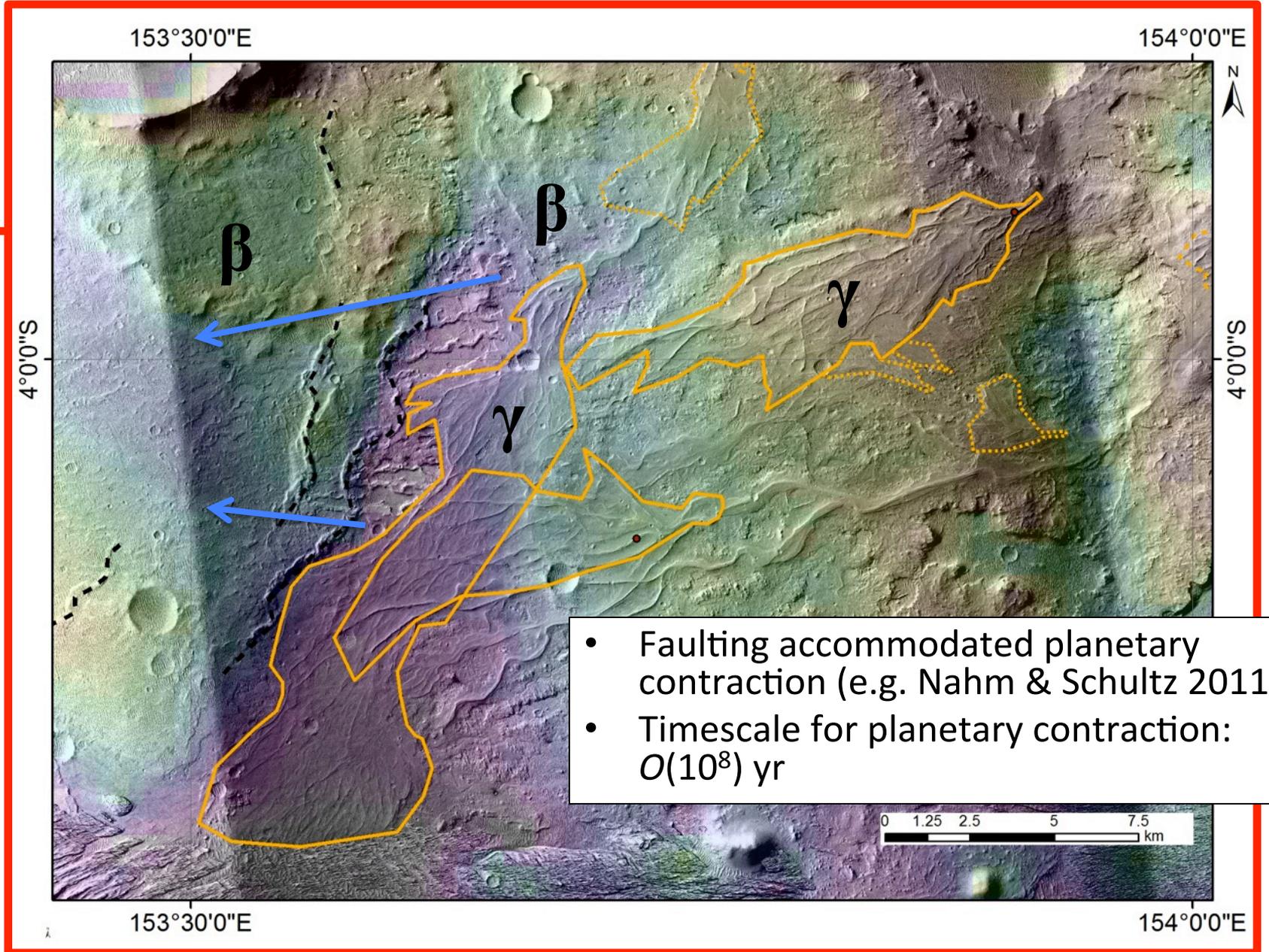
Sequence boundary example 2/3: **Thrust faulting postdates river deposits (β) and predates alluvial fans (γ)** - *At least in E of Aeolis Dorsa*

δ

γ

β

α

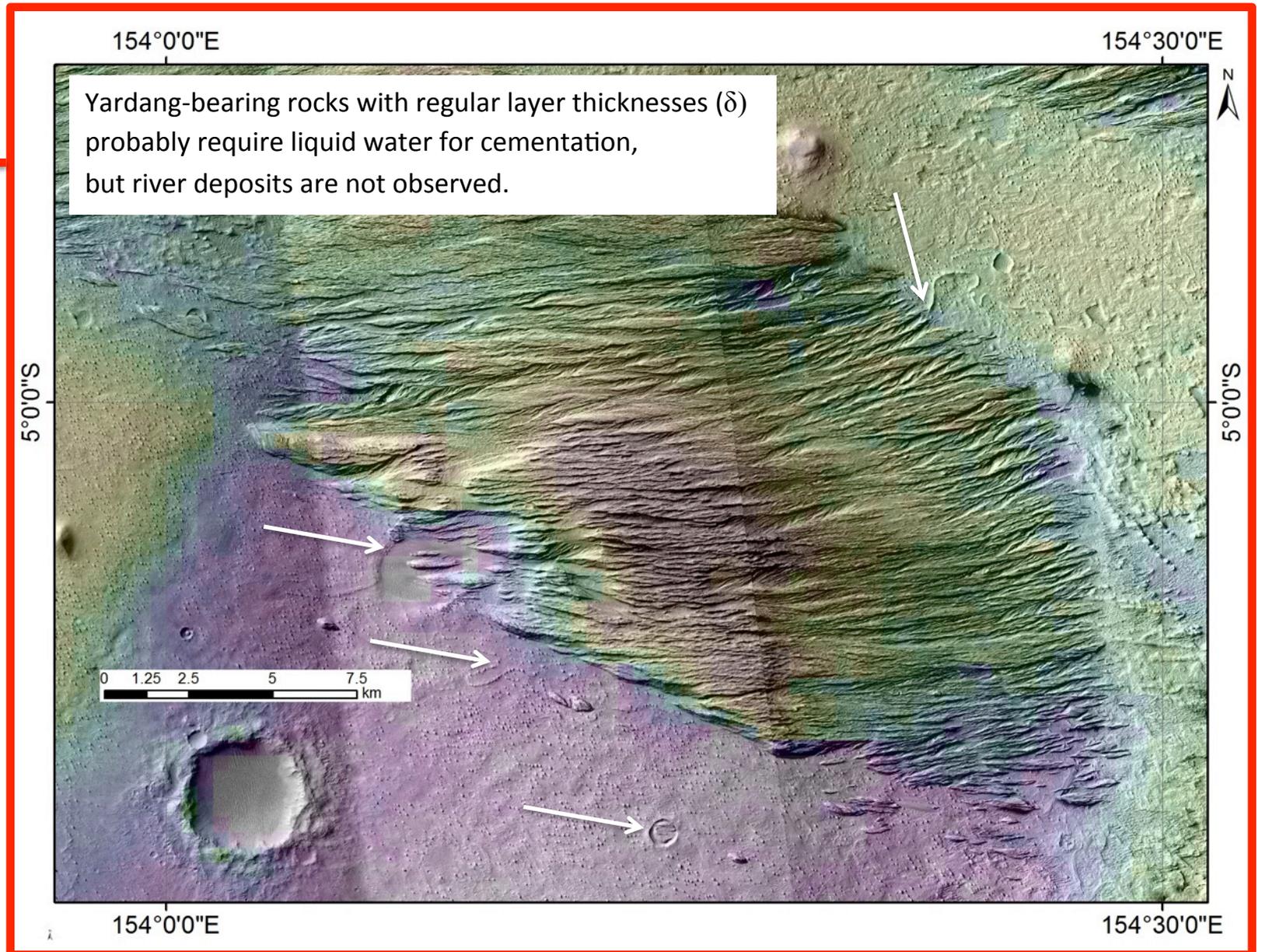


Sequence boundary example 3/3: Densely cratered surface separates river deposits (β) from yardang-bearing materials (δ)

δ

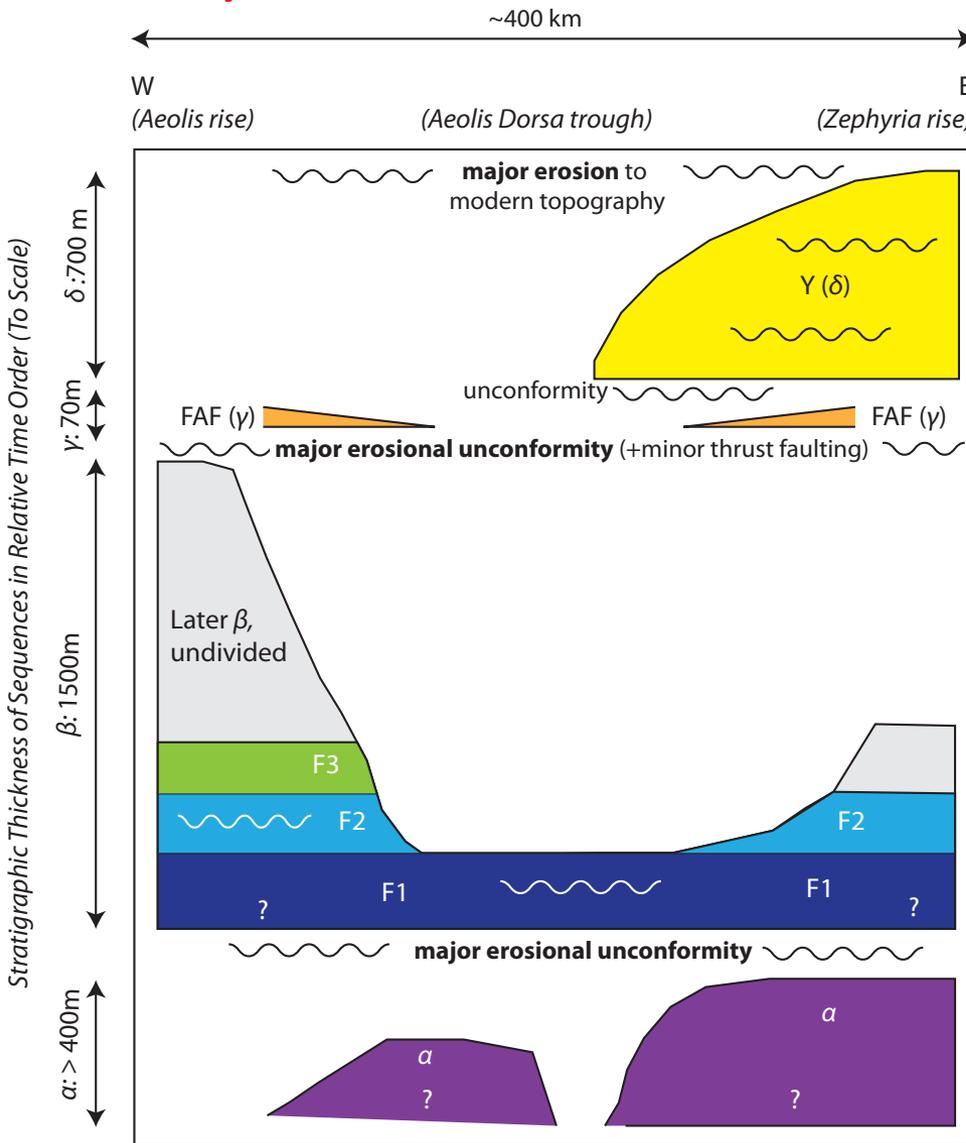
β

α



Climate proxies show that the great drying of Mars was not steady

Preliminary



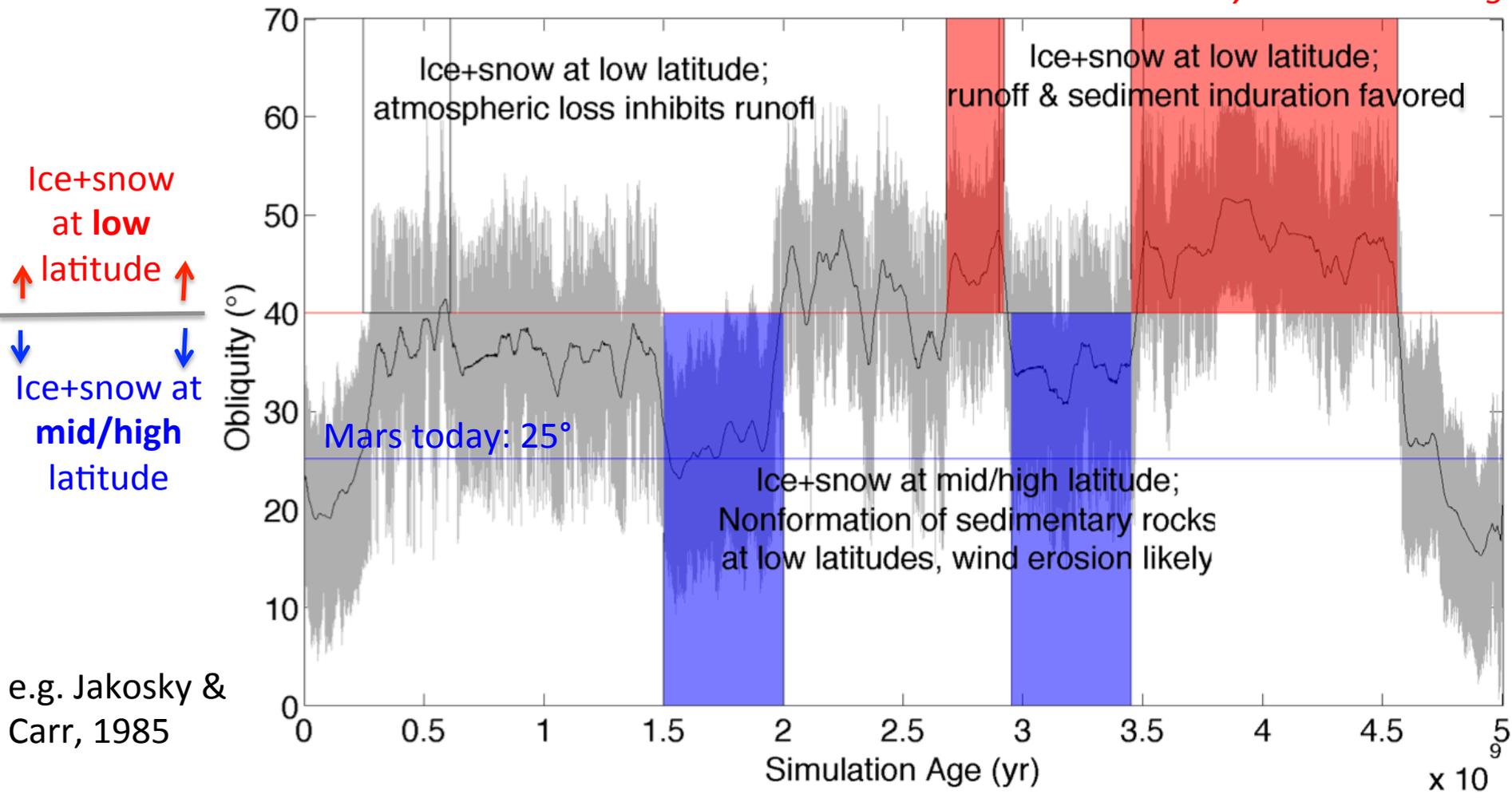
Epoch	Max. river network length (km)	Mean channel/river deposit width (m)	Interpretation: Climate conditions
Deflation to modern topography	Wind erosion, dry conditions		Dry
δ	Aqueous cementation inferred. River deposits not found.		Damp
Sub- δ unconformity	Wind erosion, dry conditions		Dry
FAF	40	30	Wet
Sub- γ unconformity	Major planar erosion Wind erosion Possible minor river erosion		Dry
Later β	Gap in fluvial-deposit record		?
F3	60	60	Very wet
F2	20	40	Less wet
F1	>500	100	Very wet
Sub- β unconformity	Major dissectional erosion Possible fluvial or glacial		?
α	River deposits not found.		?

Next: quantify using ISEE-Mars framework - Kite et al., "Seasonal melting ...," Icarus 2013

Mean-obliquity shifts are a plausible **driving mechanism** of observed shifts between erosion and deposition.

- Plate tectonics & eustasy unlikely
- Implies signal should be **global**

Shown: one of many possible Mars obliquity histories.
Calculation by John Armstrong



Conclusions

- Stratigraphy consistent with Aeolis Dorsa recording an **unusually complete** history of Early Mars fluvial environments.
(candidate global reference section)

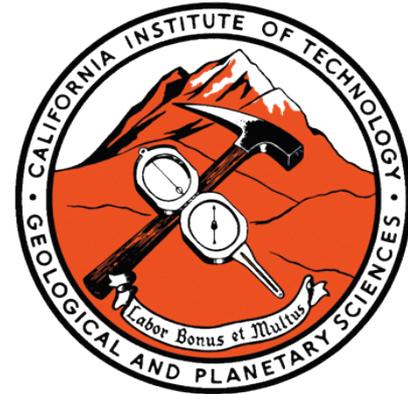
- Drying-out of Mars climate was **not steady**.

Goal: extract enough constraints to resolve disagreements about the cause and duration of warm/wet climates.

Requires quantitative models linking sed./strat. to climate.

More information: Kite et al., 'Seasonal melting ...' *Icarus* 2013, ←
Kite et al., 'Growth and form ...' *Geology* 2013, ←
Kite et al., 'Pacing fluvial ...' *Icarus* 2013,
Kite et al., 'Paleopressure ...' *Nature Geoscience*, 'accepted in principle,' arxiv:1304.4043
gps.caltech.edu/~kite

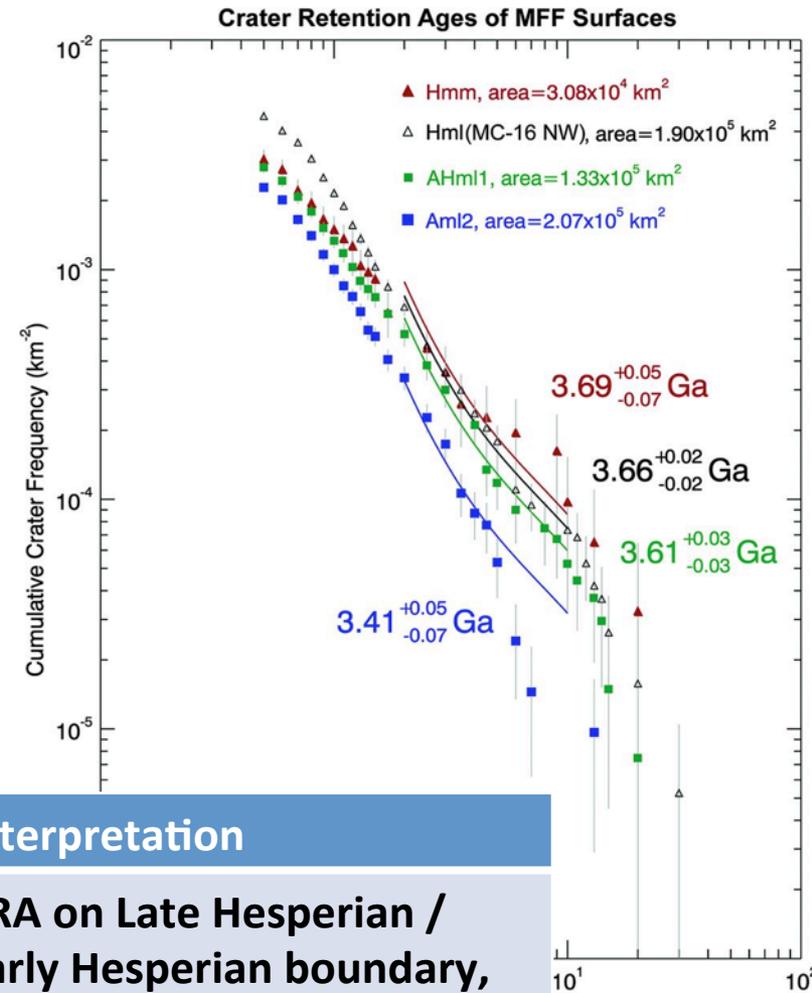
With thanks to: Mike Lamb, Woody Fischer, Devon Burr, Robert Jacobsen, Rebecca Williams, Roman diBiase, Or Bialik, and all the participants in Caltech's Mars Fluvial Geomorphology Reading Group



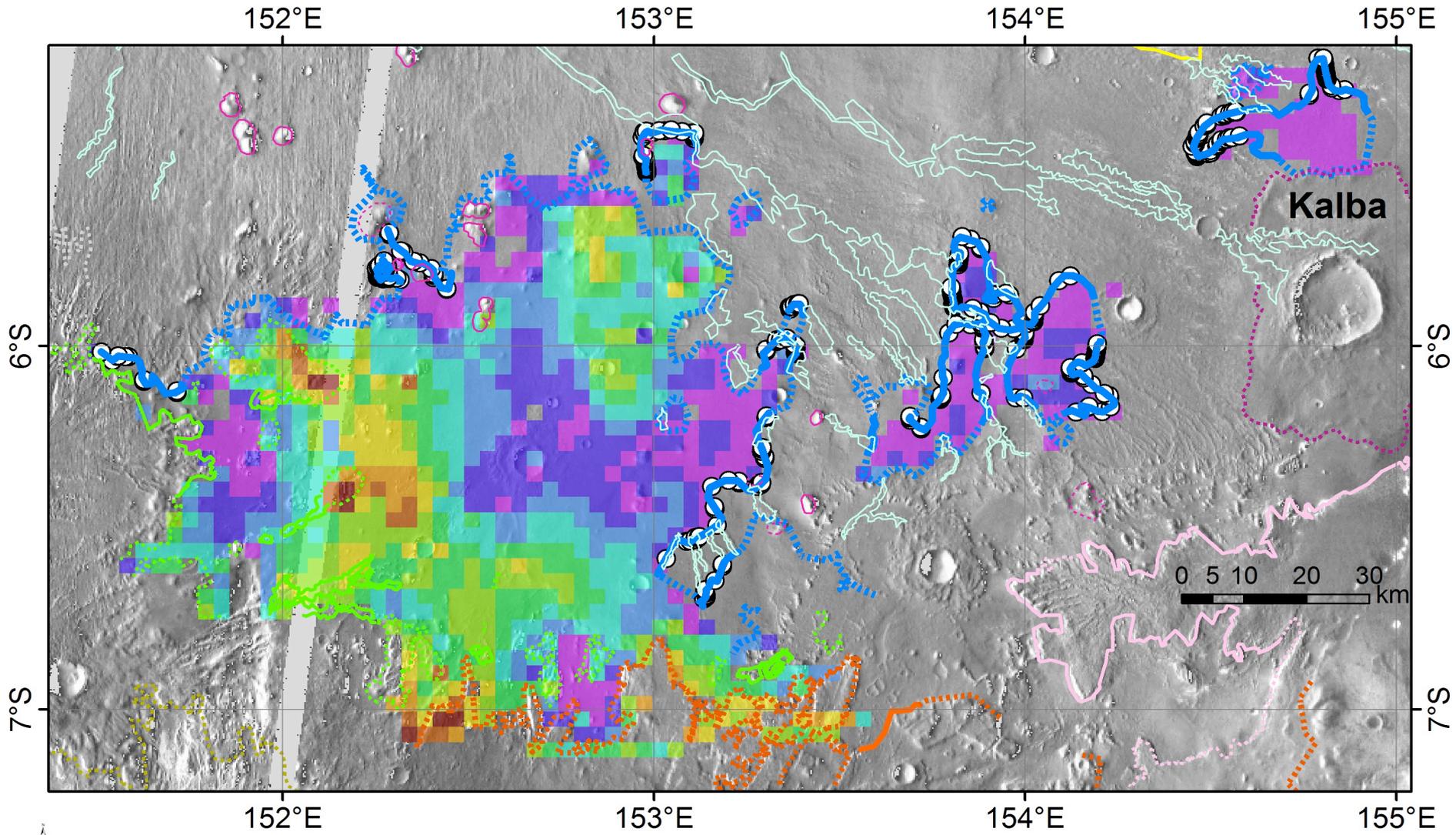
End of
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Crater density consistent with Noachian/Hesperian boundary – and density of rivers suggests that

- Major Amazonian erosion has occurred (mesas)
- Zimbelman & Scheidt (Science), 2012 “craters on a nearby exposure of middle-member material (**superposed on the [rivers]**) indicate a late Hesperian age
- CRA of rivers is on Late Hesperian/ Early Amazonian boundary

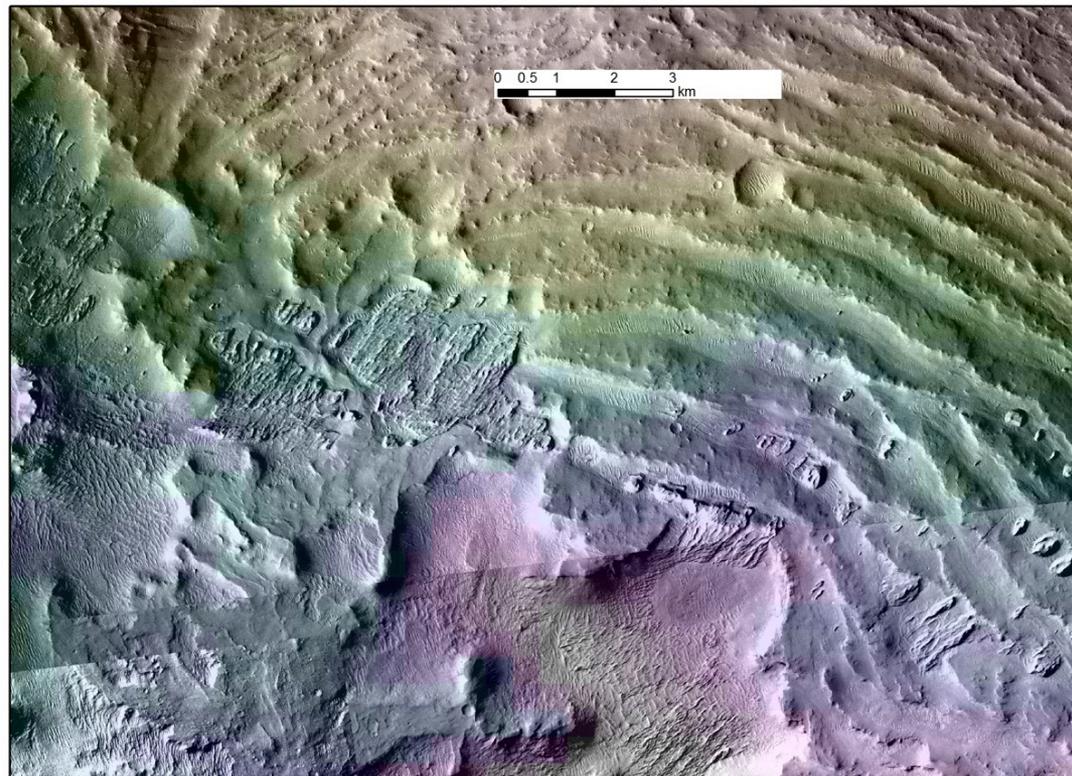


	N(1)	N(2)	N(5)	Interpretation
Postfluvial + Undetermined	2049±158	634±88	37±21	CRA on Late Hesperian / Early Hesperian boundary, consistent with overall correlation to Noachian/Hesperian boundary
Postfluvial only	1049±113	415±71	37±21	
Synfluvial				



Evidence for additional unconformities

- Some evidence for unconformity separating alluvial fans and yardang-bearing materials
- Strong evidence for unconformities *within* fluvial deposits, at multiple scales

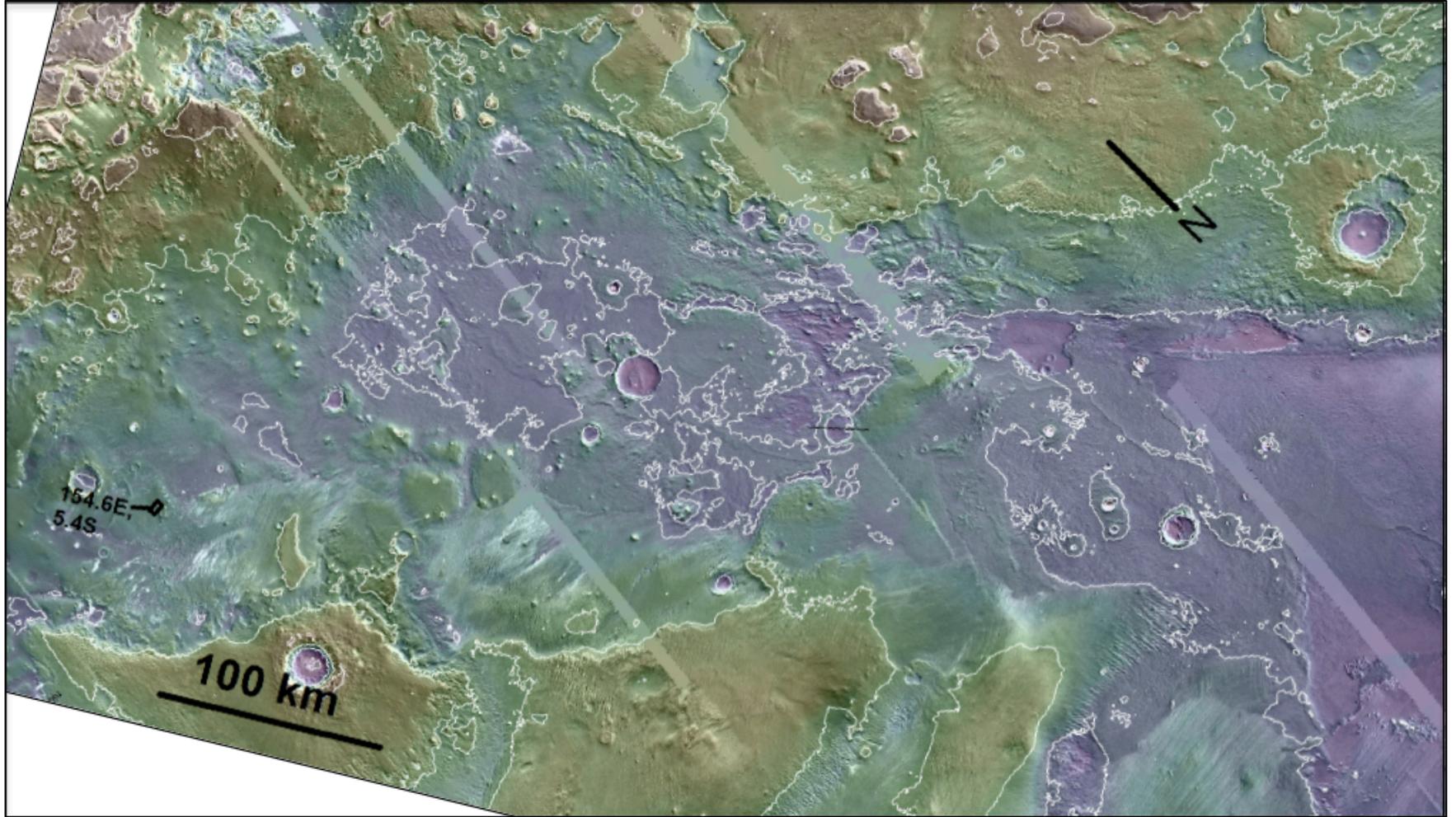


δ

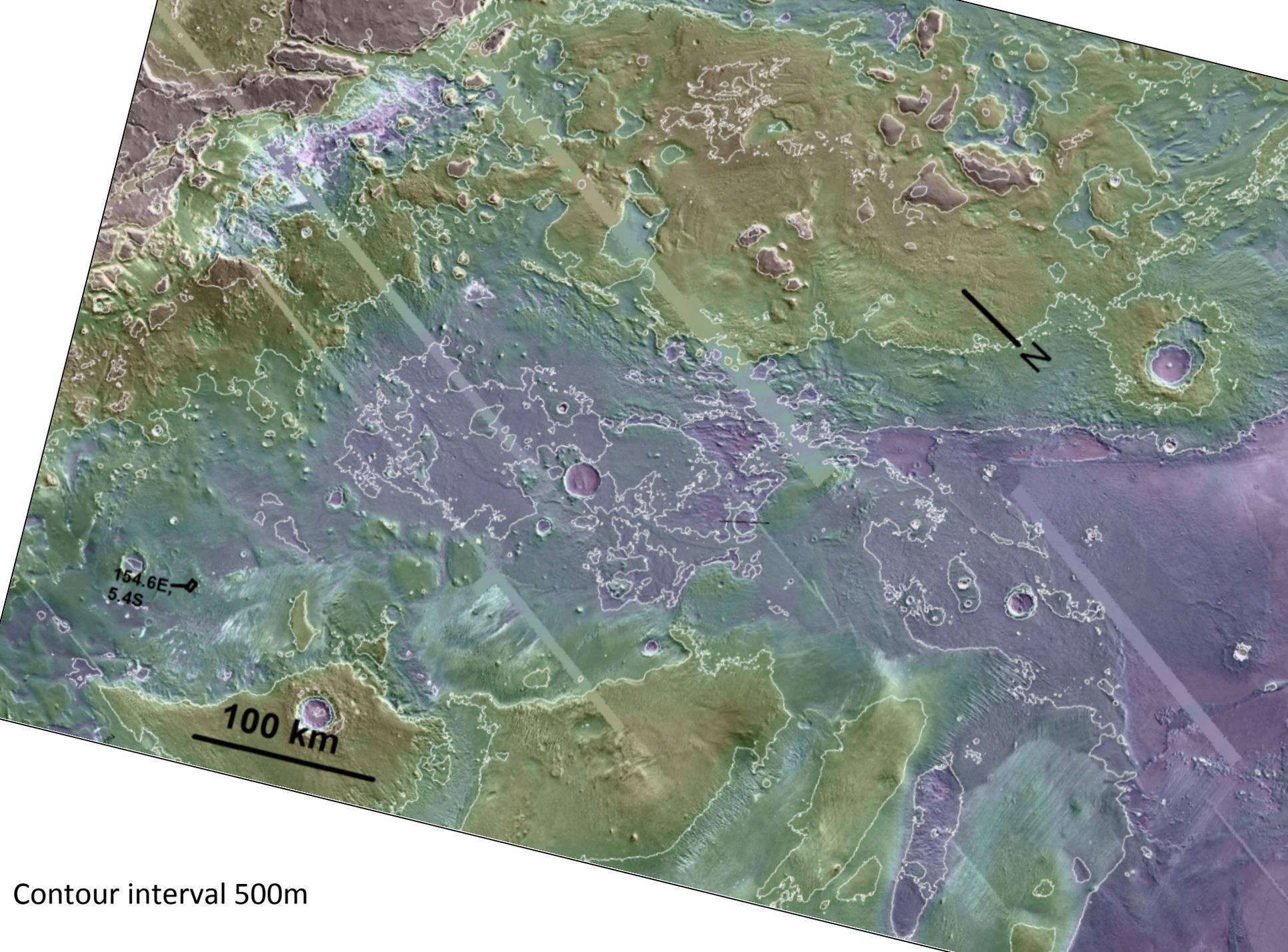
γ

β

α



Contour interval 500m



164.6E, 5.4S

100 km

Contour interval 500m